



**SOHO 30 years:**

**New and Future missions building on SOHO's legacy**

Anik De Groof  
ESA mission manager for SOHO and Solar Orbiter  
Solar Physicist inspired by SOHO

SOHO 30 years celebration, Orsay, 10/12/2025



# SOHO's legacy



Over the past 30 years, SOHO made a profound and lasting impact on the field of solar and space physics:

- groundbreaking **discoveries**, shaping the direction of heliophysics research
- inspiring a whole generation of **solar physicists** (~300 PhD theses)
- influencing the design, instrumentation, and scientific objectives of subsequent **space missions**

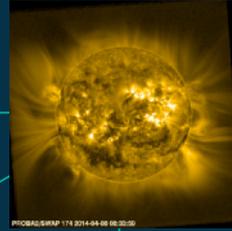
SOHO taught us the importance of:

- coordinated multi-instrument observations
- long-term solar monitoring
- continuous, stable observations for space weather monitoring and forecasting

And it also showed the holes in our understanding and how to potentially address them!

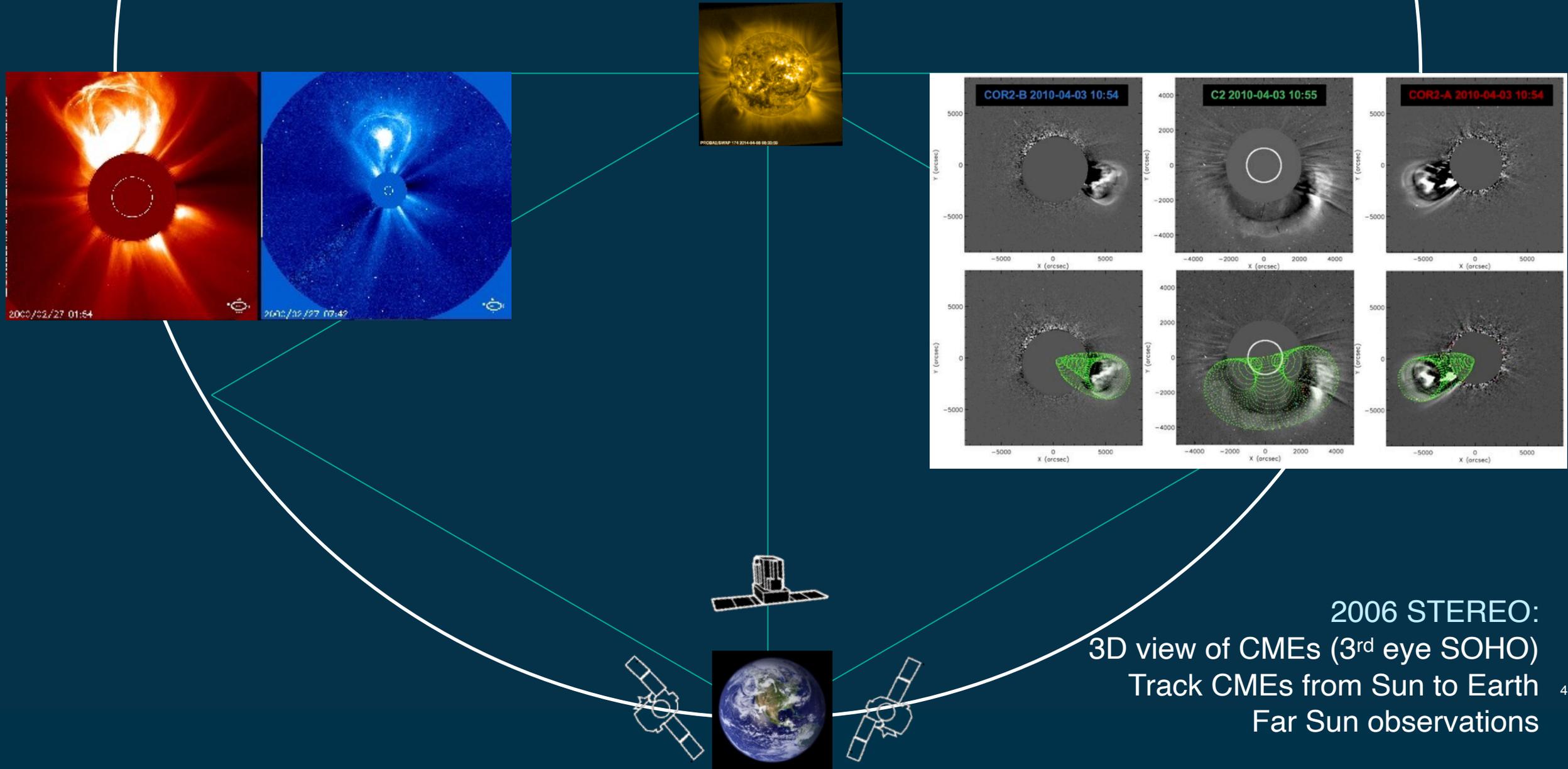


# New ways of looking at the Sun



1995 SOHO (L1):  
First 24/7 view on flares, CMEs  
Space weather awareness  
Long-term monitor for  
Helioseismology

# New ways of looking at the Sun – stereoscopic view

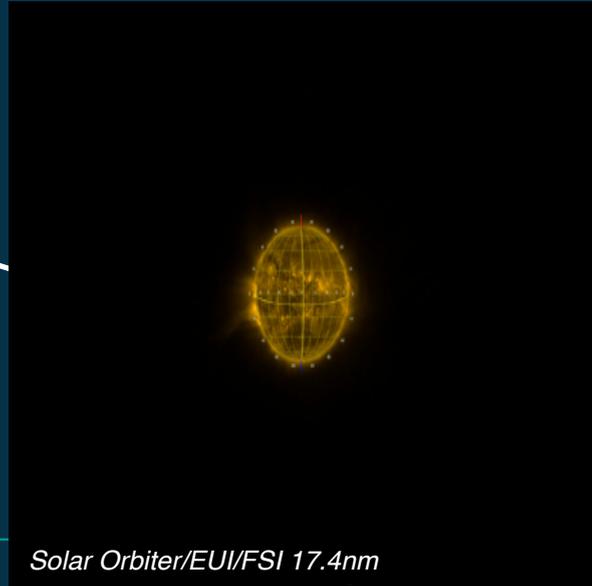
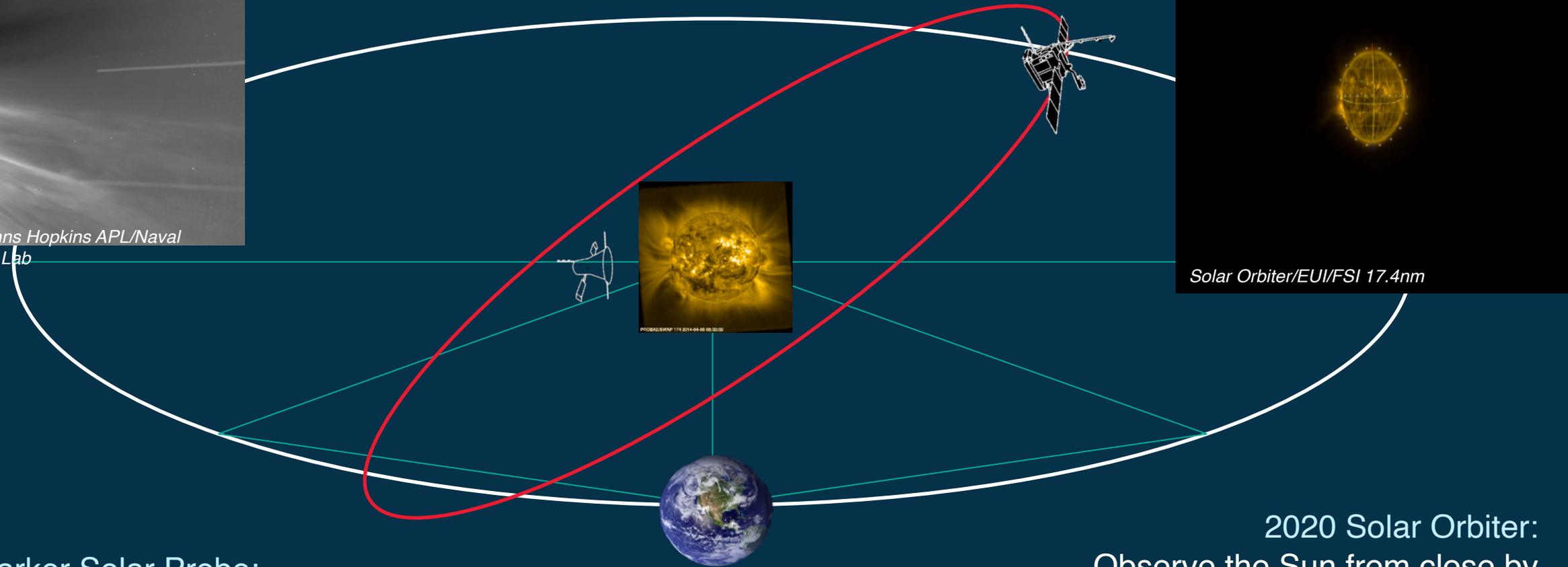


2006 STEREO:  
3D view of CMEs (3<sup>rd</sup> eye SOHO)  
Track CMEs from Sun to Earth  
Far Sun observations

# New ways of looking at the Sun – Close in & polar view



NASA/Johns Hopkins APL/Naval Research Lab



2018 Parker Solar Probe:  
Dive close to the Sun to study how  
corona becomes solar wind (in situ)

2020 Solar Orbiter:  
Observe the Sun from close by  
Move out of the ecliptic: poles & latitude scans  
Combine Remote-Sensing with In-Situ

Far Sun observations – multi-spacecraft observations

# Newborn missions with innovative designs: Proba-3

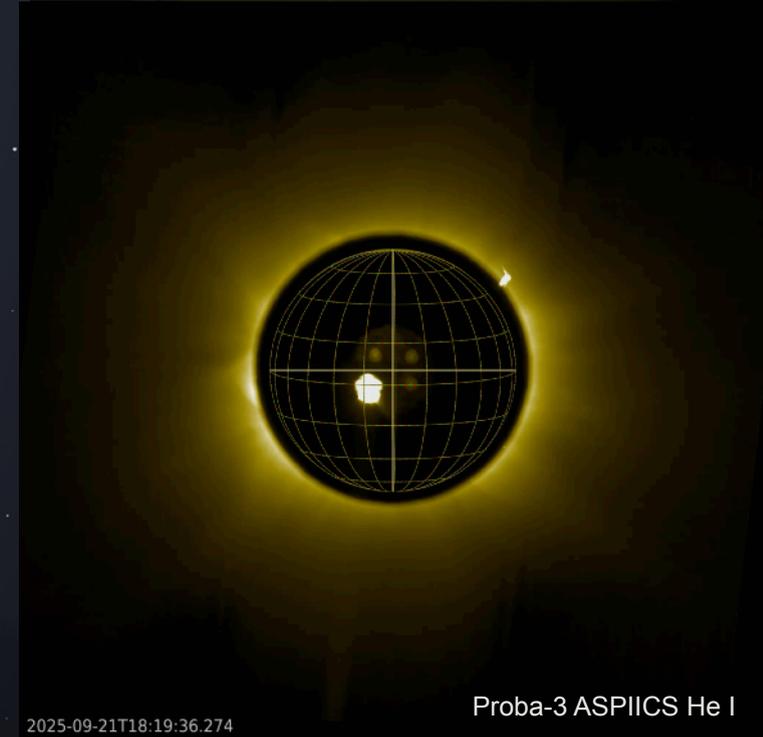


Dec 2024 Proba-3:

World's first precision formation flying mission  
Coronagraphy closer to the solar disk than ever



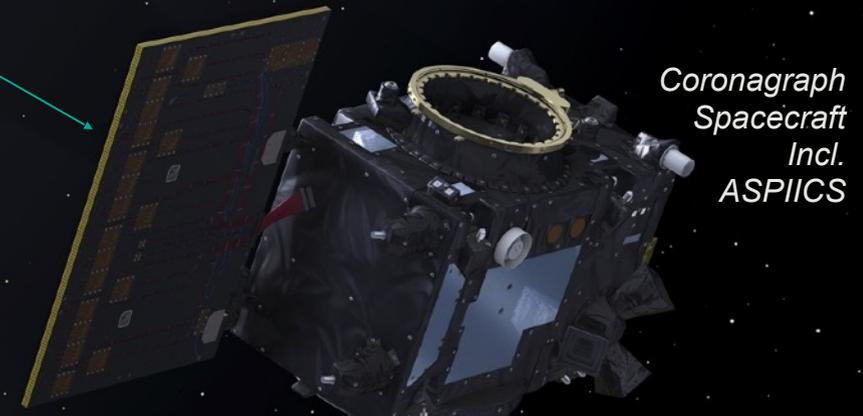
Occulter  
Spacecraft



Proba-3 ASPIICS He I

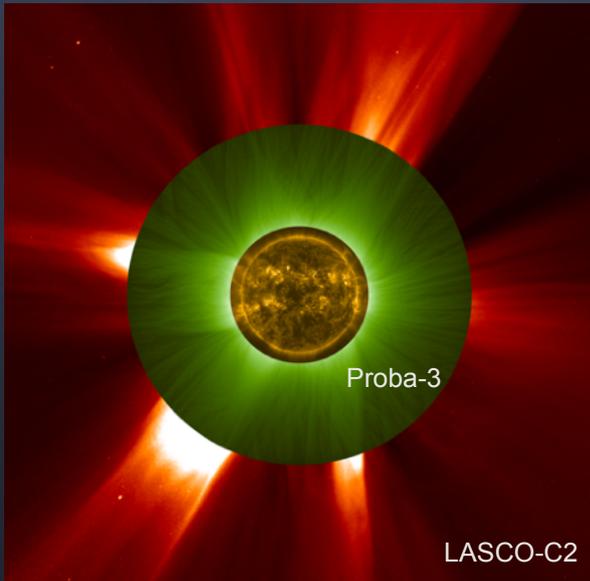
2025-09-21T18:19:36.274

150 meters



Coronagraph  
Spacecraft  
Incl.  
ASPIICS

Two spacecraft flying in precise formation  
Position accuracy need: millimetric



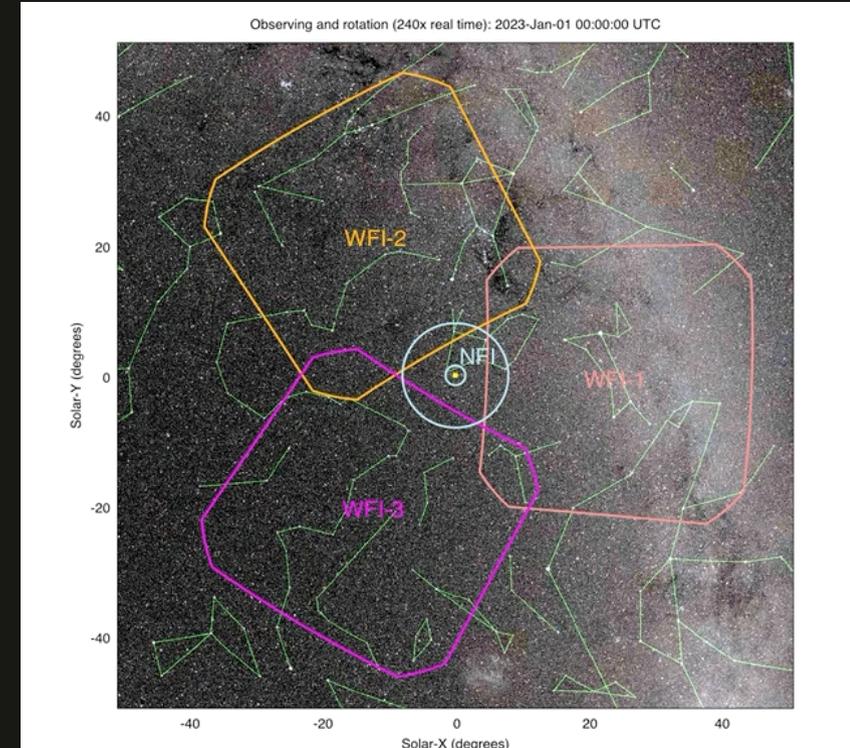
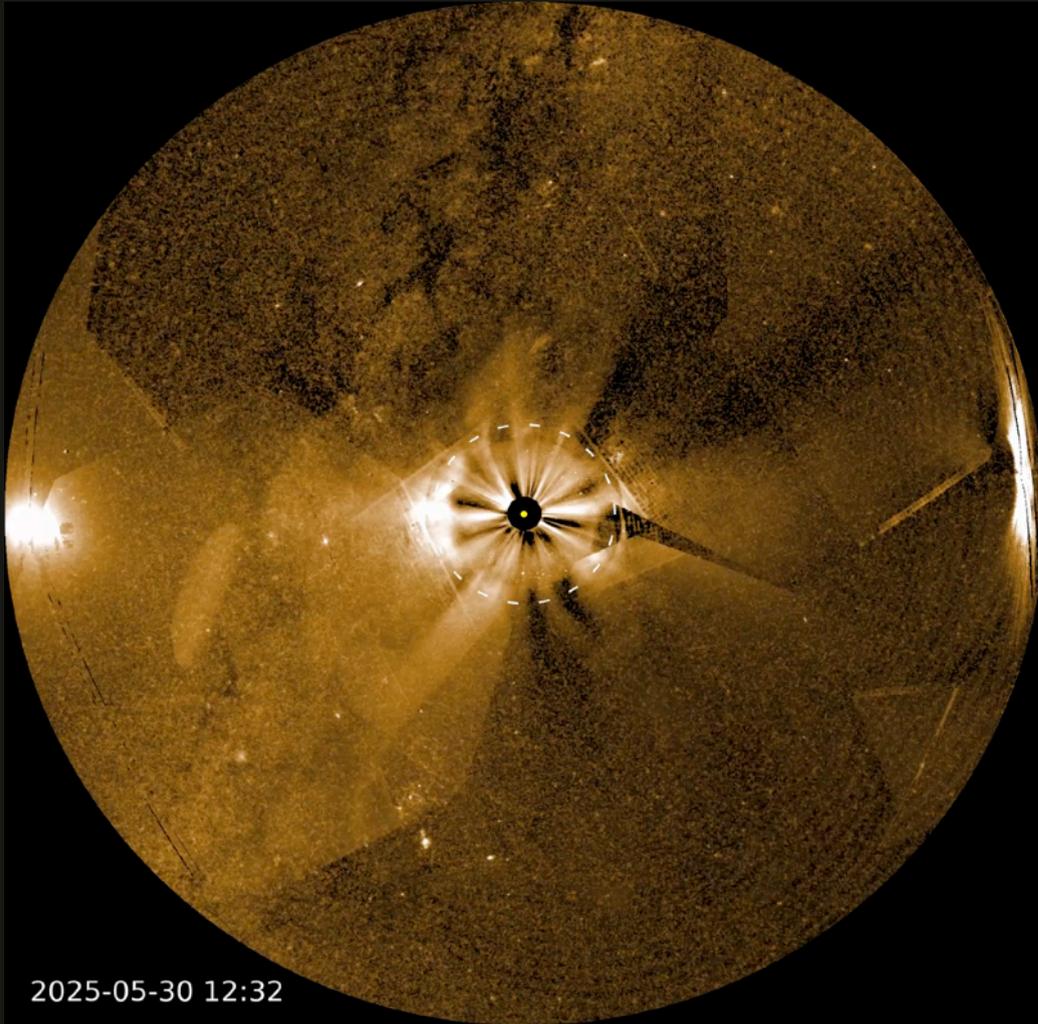
Proba-3

LASCO-C2

# Newborn missions with innovative designs: PUNCH



PI: Craig DeForest (SwRI, Boulder)



**March 2025 PUNCH:**  
Unique View of the Solar System  
Track solar transients in all directions  
Constellation of 4 spacecraft in polar LEO



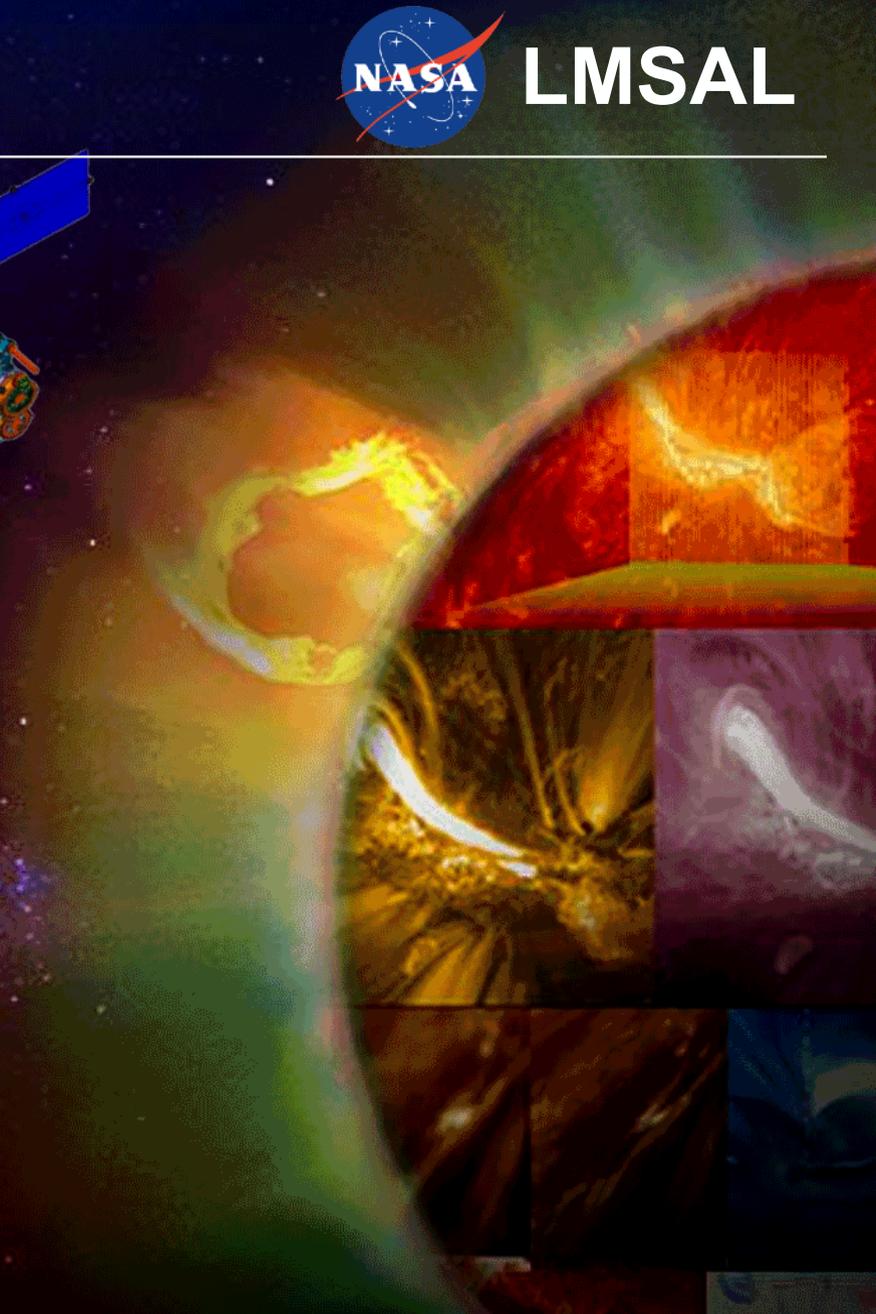
# Future Solar Missions – MUSE

Payload: MUlti-slit Solar Explorer **MUSE**

PI: Bart DePontieu (LMSAL)

Main science questions:

1. Investigate the causes of coronal heating and instability (flares and CMEs)
2. Gain insight into the basic plasma properties of the corona



# Future Solar Missions – MUSE

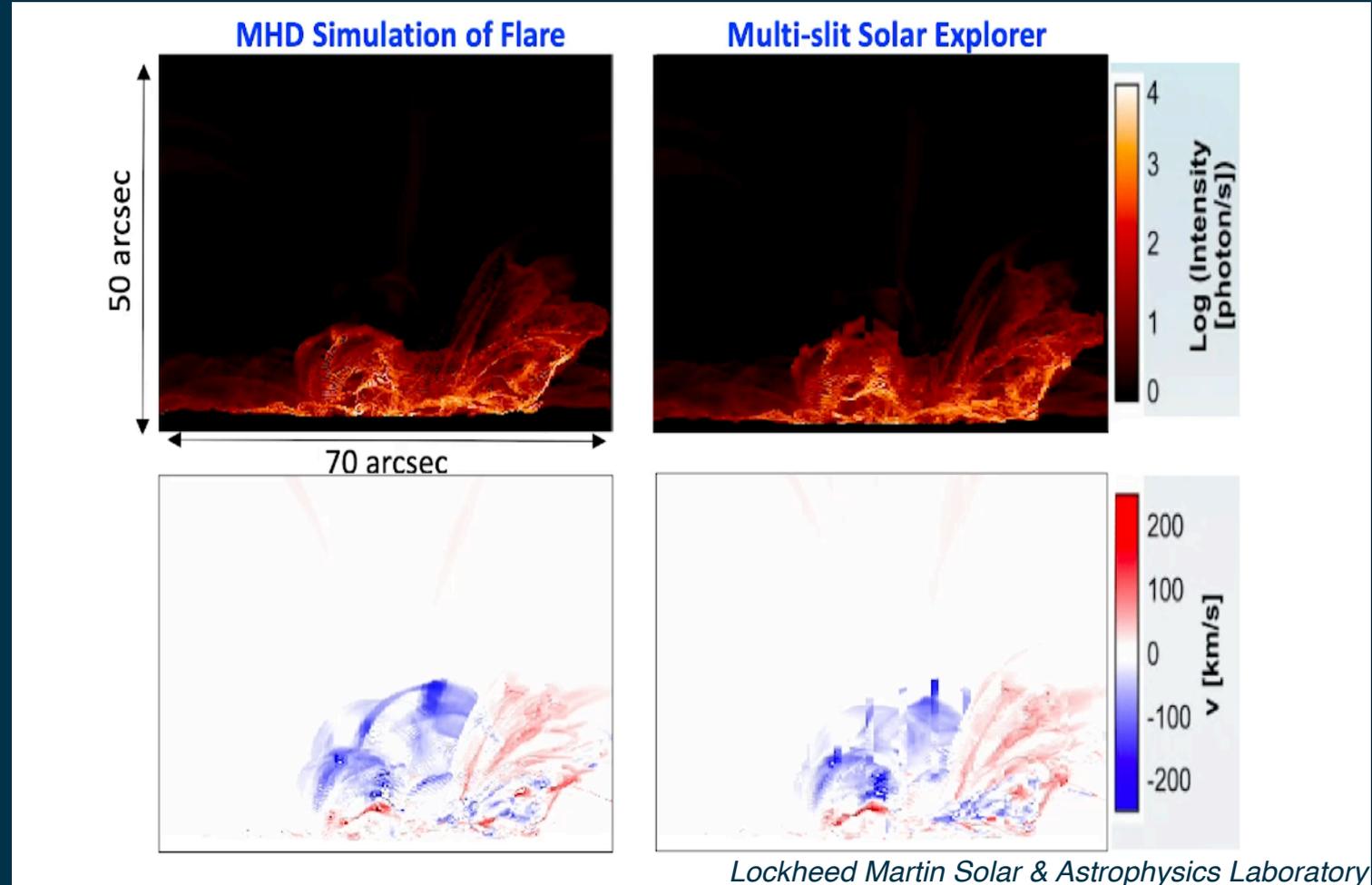
## How?

EUV spectral imaging of corona  
new imaging spectroscopy techniques  
observe radial motion and heating at  
10x current resolution and 35x faster

## Launch and orbit:

Launch expected early 2027

In May '25, MUSE passed CDR



**Spectrograph** will raster 35 slits across same scene in 3 EUV bands  
**Context Imager** sampling at 304 Å and 195 Å for spatial context

# Future Solar Missions – Solar-C



## Payload

- EUV High-throughput Spectroscopic Telescope **EUVST**:
  - EUV slit spectrographs: Long-WL and Short-WL camera (~IRIS & Hinode/EIS resp.)
  - Slit-jaw imager for context (~IRIS)
- Solar Spectral Irradiance Monitor **SoSpIM**:
  - solar irradiance in EUV and Ly alpha (<1s cadence)

## Main science questions:

- How are the hot corona and solar wind generated?
- When and how do Solar Flares occur?

## How to address those:

- Seamless temperature coverage from 10,000 K to 15 million K
- High spatial (0.4") and temporal (1 sec) resolution
- Velocity resolution of 2 km/s

Very complementary to MUSE



Launch and orbit:  
Launch scheduled ~2030  
Orbit: Sun-synchronous LEO



# Future Solar Wind Missions – SMILE

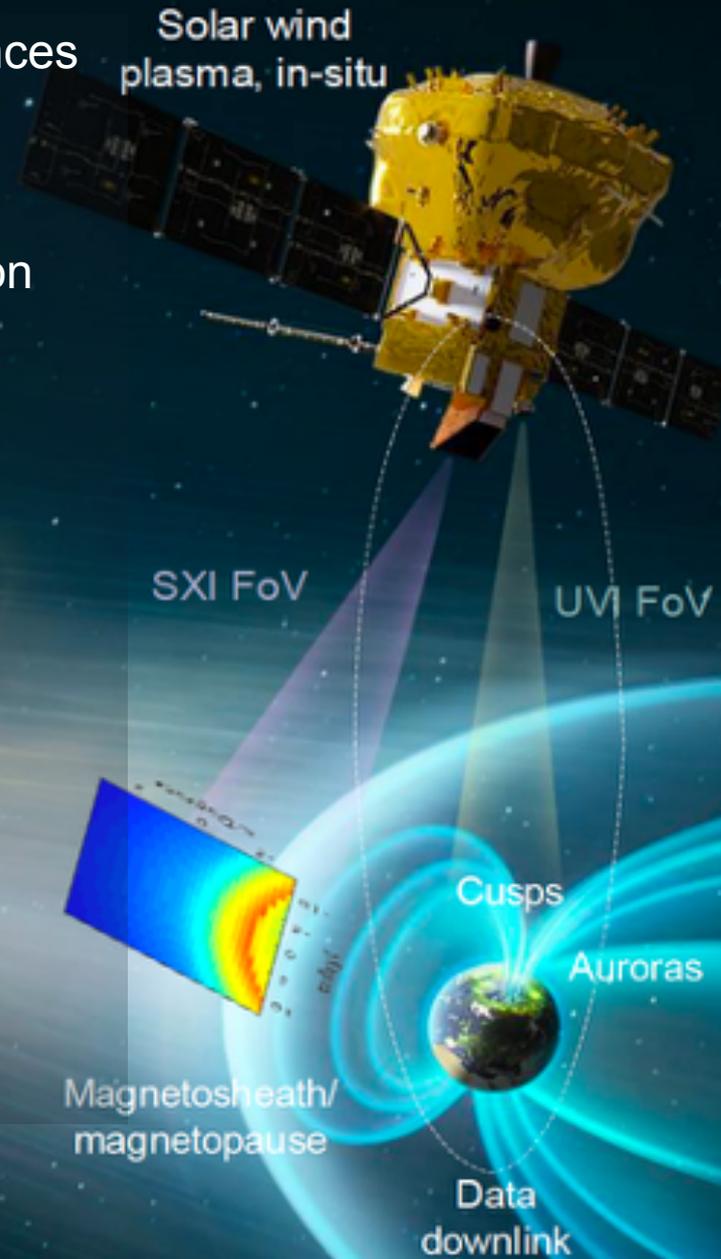
Collaboration ESA & Chinese Academy of Sciences

## Main science goal:

complete understanding of Sun–Earth connection by measuring the solar wind and its dynamic interaction with Earth’s magnetosphere.

## How to address those:

- Measure & image solar wind and magnetosheath plasma
- Observe cusp, magnetopause and magnetosheath for first time in soft X-rays
- Resume auroral imaging after termination IMAGE and Polar operations 15 yrs ago



Payload:

**LIA** Light Ion Analyser  
5 eV-20 keV  
4  $\pi$  FOV

**MAG** flux-gate magnetometer  
B measured up to 40 Hz

**SXI** X-ray imager  
wide FOV 16 x 27°

**UVI** UV imager  
155-175 nm, 10° FOV

# Future Solar Wind Missions – SMILE

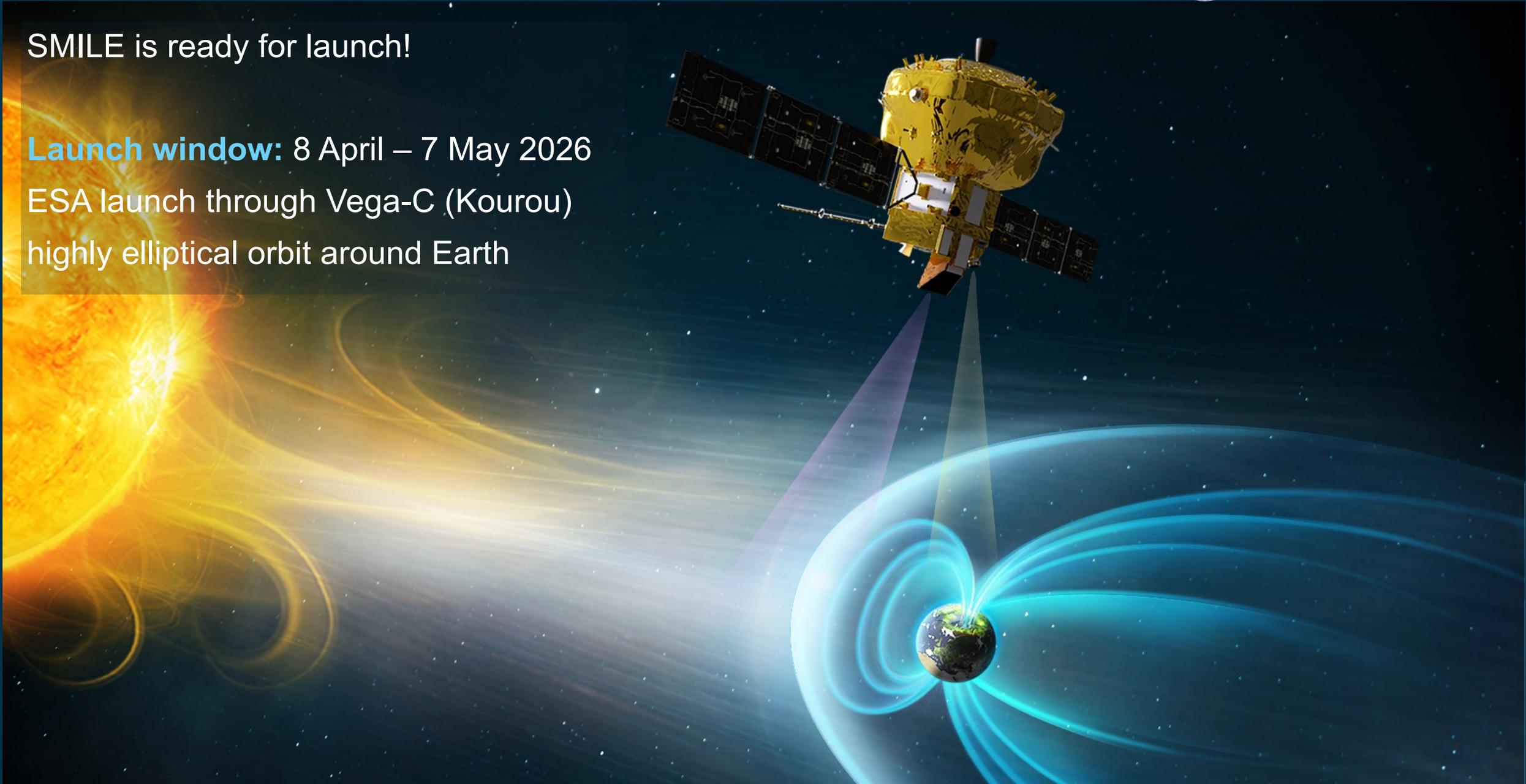


SMILE is ready for launch!

**Launch window:** 8 April – 7 May 2026

ESA launch through Vega-C (Kourou)

highly elliptical orbit around Earth



# Future Space Weather missions – SWFO-L1



SWFO-L1 = Space Weather Follow On - Lagrange 1

NOAA/NASA mission

Coronagraph CCOR-2 is considered the successor of LASCO for space weather

**Launched** September 24, 2025

Arrives at L1 in January 2026, expects to be operational mid 2026.

Full data product validation expected in 2027, after cross-calibration with LASCO

Coronagraph data will be on Earth within 30 mins

In situ data will be available within 5 minutes

- Solar Wind Plasma Sensor (SWiPS),
- SupraThermal Ion Sensor (STIS),
- Magnetometer (MAG)



# Future Space Weather missions – Vigil

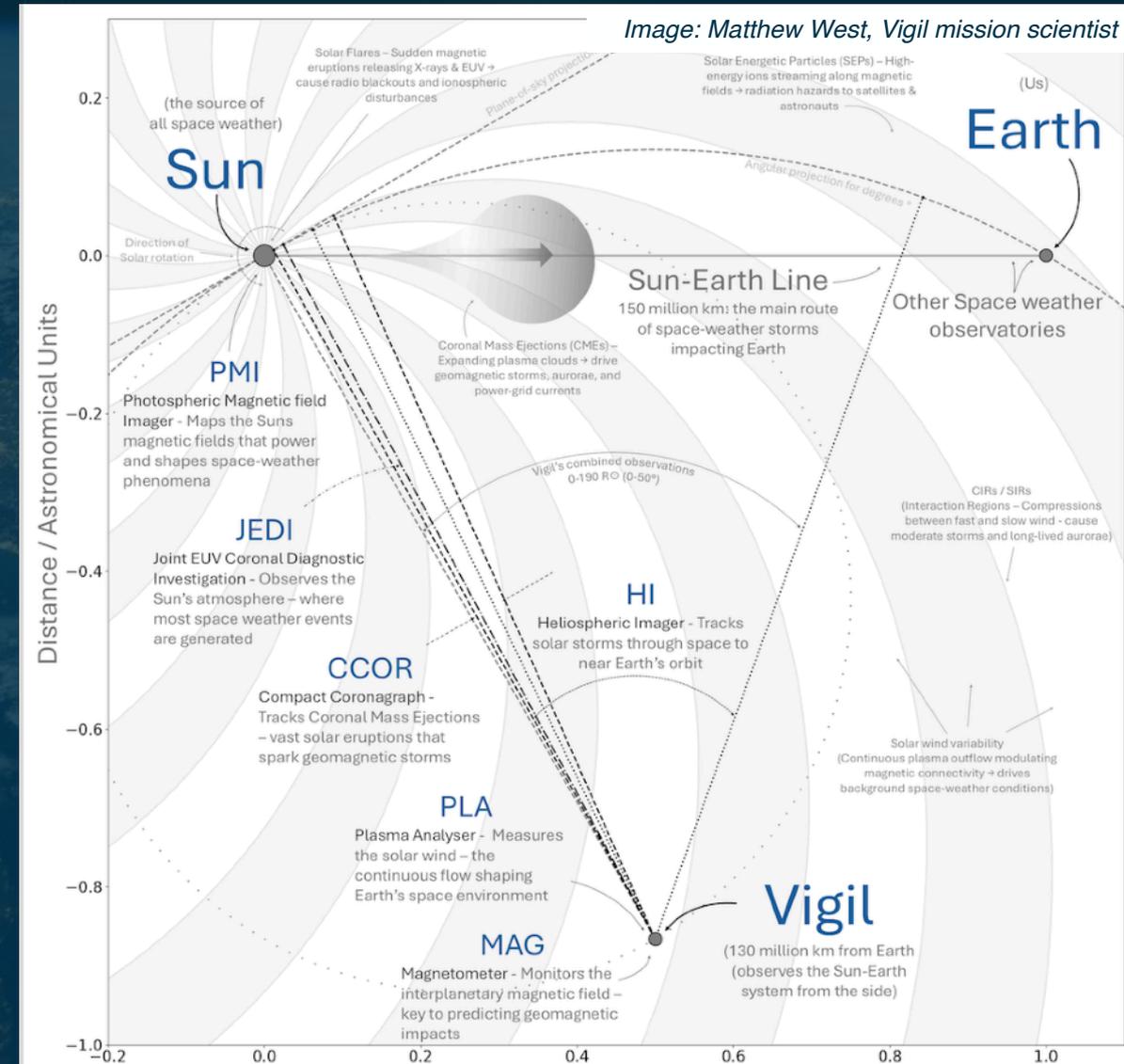


## ESA space weather mission at L5

- Early warning of emerging solar weather conditions
- Forecast up to 4/5 days in advance
- Nowcasting through LowLatency data
- Improved predictions of CME impact time & location

## Payload

- Compact coronagraph **CCOR-3** 
- Heliospheric Imager **HI**  
- Photospheric Magnetic field Imager **PMI** 
- EUV Imager **JEDI** 
- Plasma Analyser **PLA** 
- Magnetometer **MAG**  



# Future Space Weather missions – Vigil

## Mission timeline

- Launch scheduled for **Q3 2031**
- Transfer to L5 expected to last 3 years
- Operational phase can start at  $30^\circ$  separation from Earth

